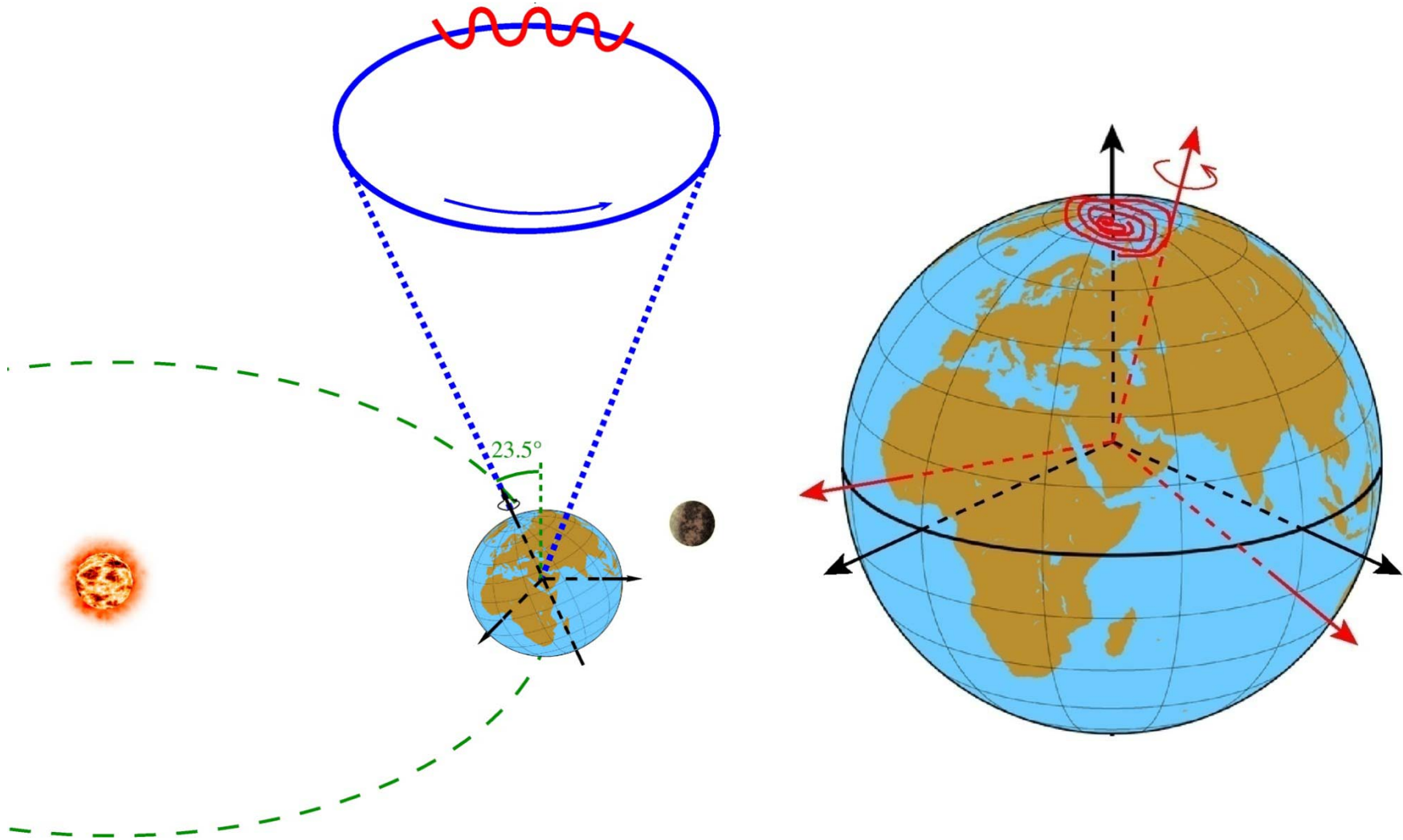


# Introduction to VLBI, IVS, VGOS, and GGOS

**Axel Nothnagel**

Institute of Geodesy and Geoinformation  
University of Bonn, Germany

- **Geodetic VLBI principles**
  - Variable Earth rotation
  - Radio sources
  - Radio telescopes
  - Data flow and analysis
- **International VLBI Service for Geodesy and Astrometry**
  - Operations
  - Observing network
- **VLBI Global Observing System**
- **Global Geodetic Observing System**



Courtesy Thomas Artz.



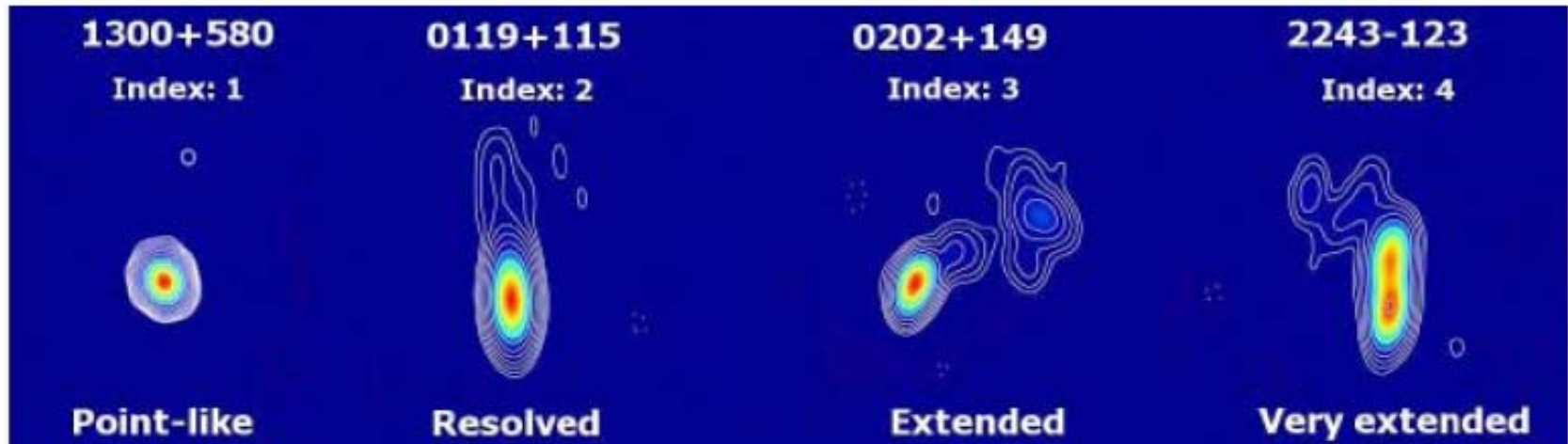
**Active galactic nuclei, galaxies, quasars**

**Distance 2 – 8 billion light years**

**Point sources**

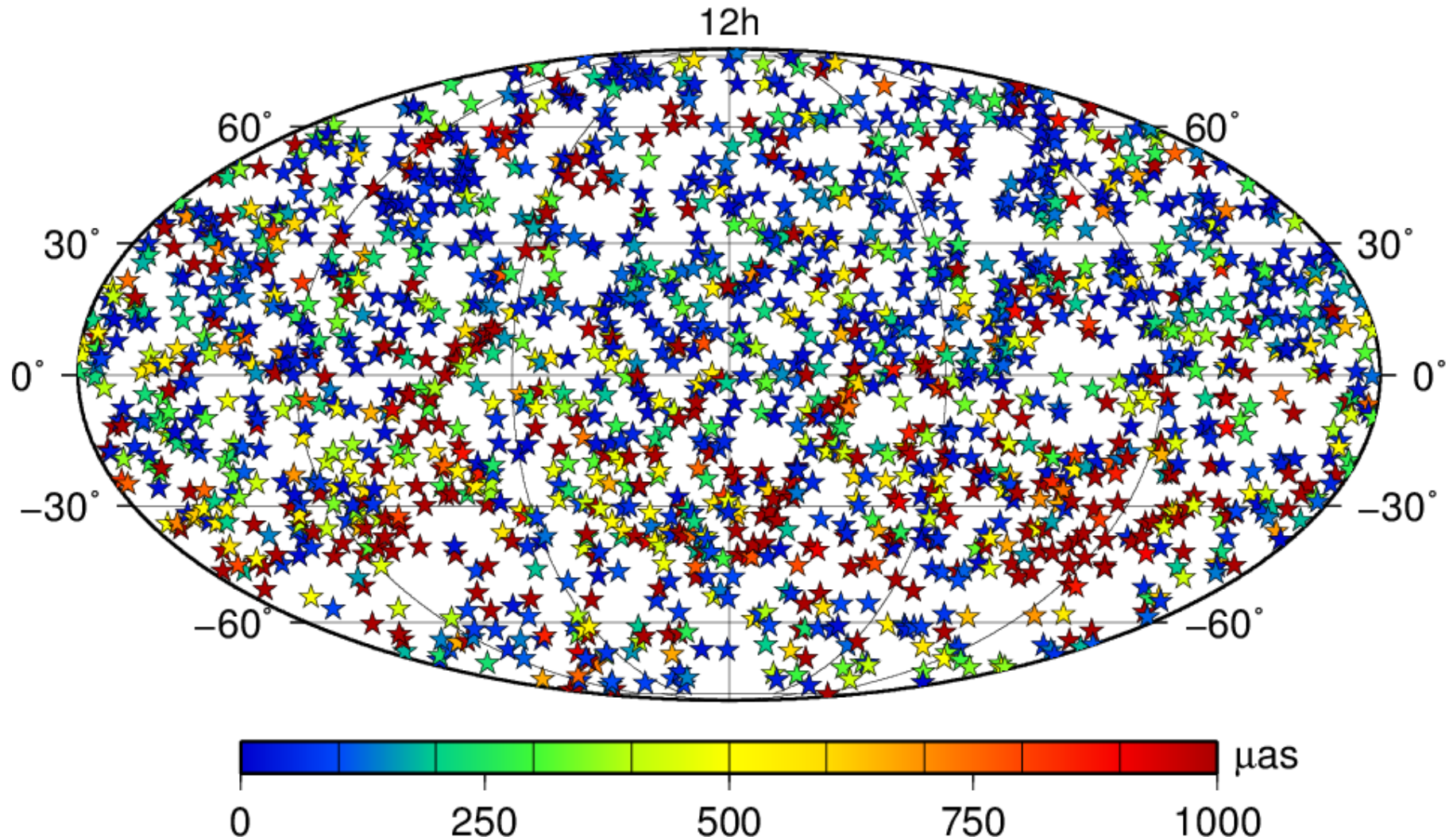
**No proper motions**

**→ quasi-inertial reference system**



Collioud and Charlot (2009)

# International Celestial Reference Frame (ICRF)





**Wetzell, Germany**



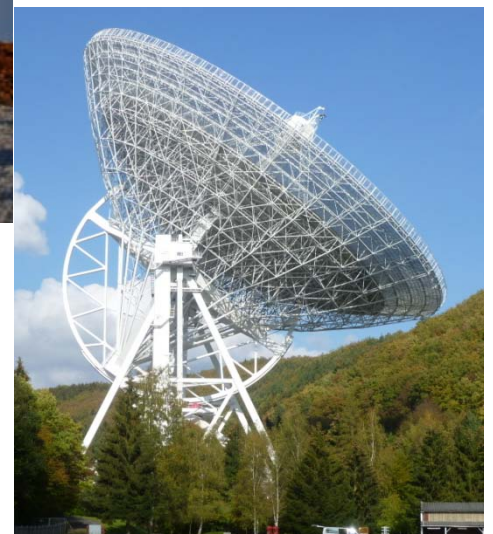
© IVS

**O'Higgins, Antarktica**

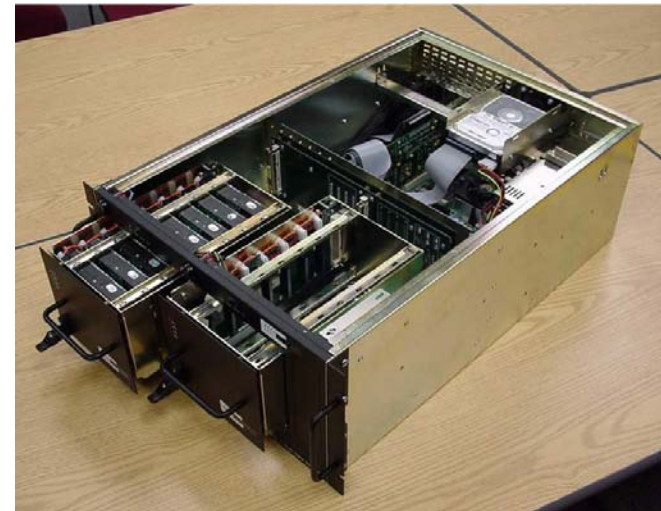
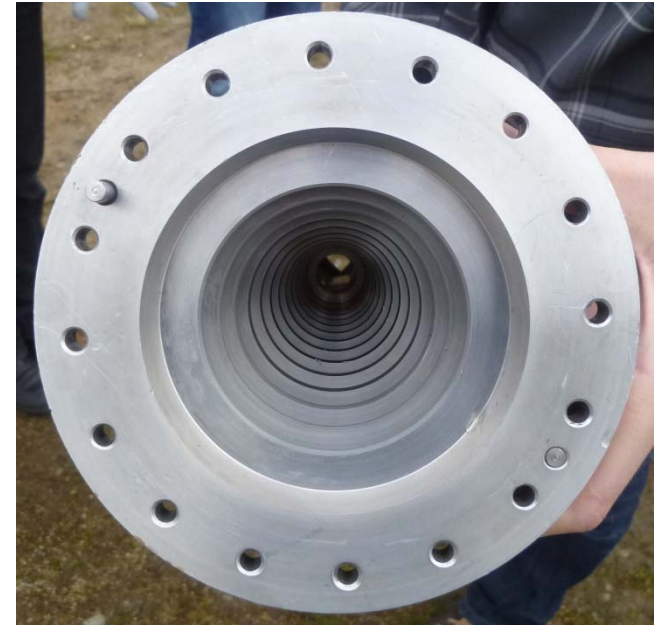
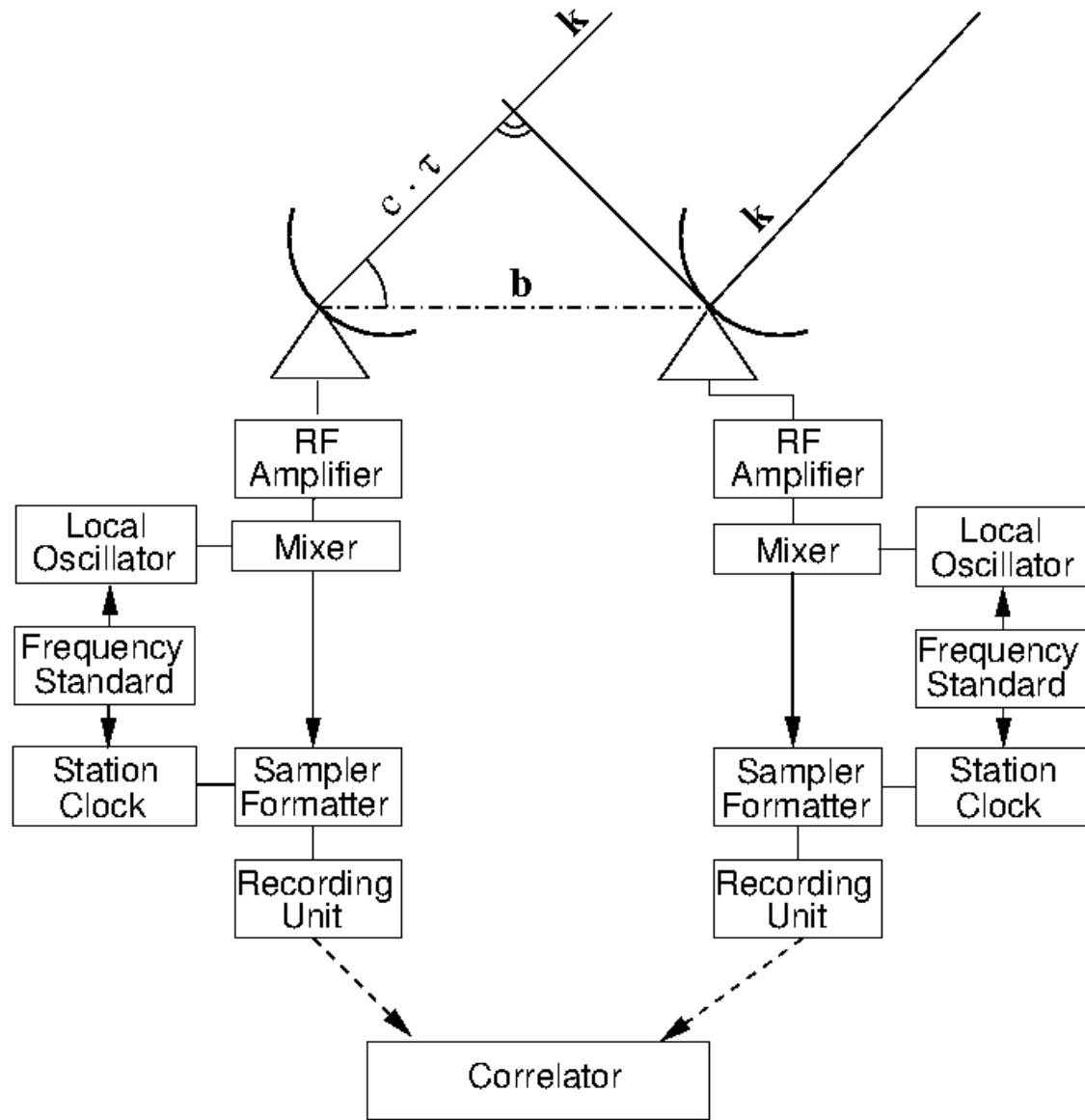
© BKG

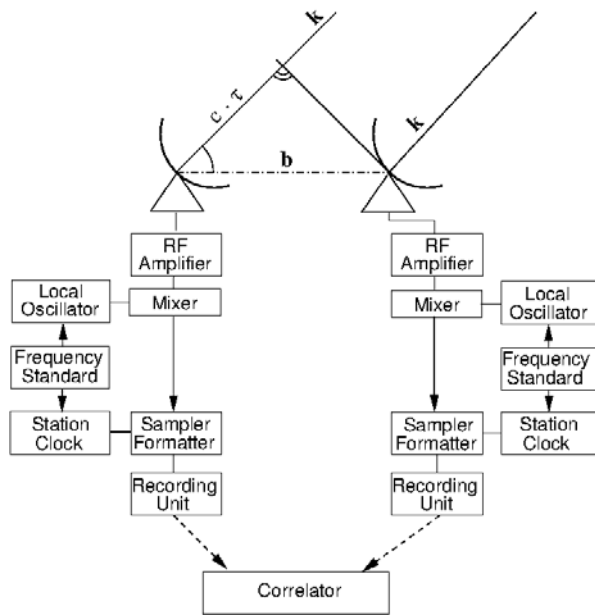


**Urumqi, China**

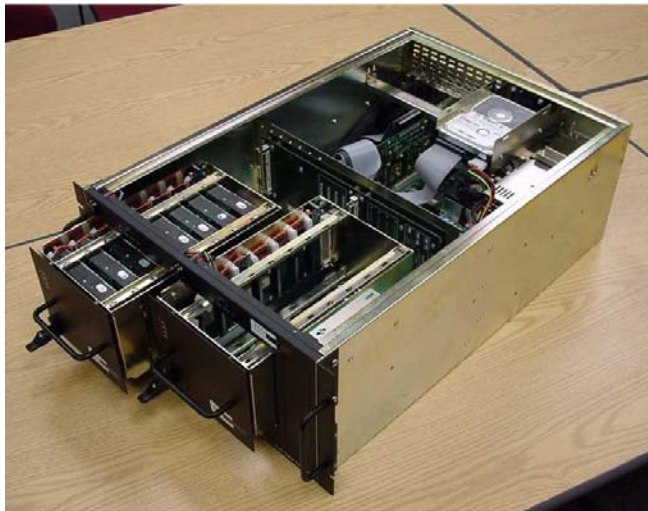


**Effelsberg, Germany**

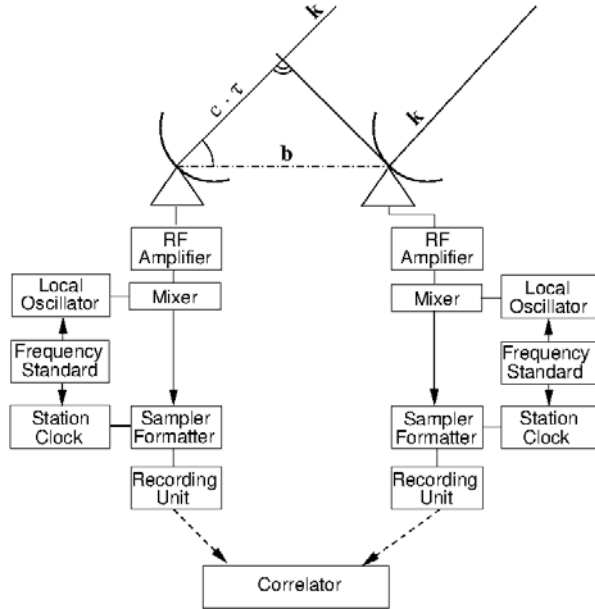




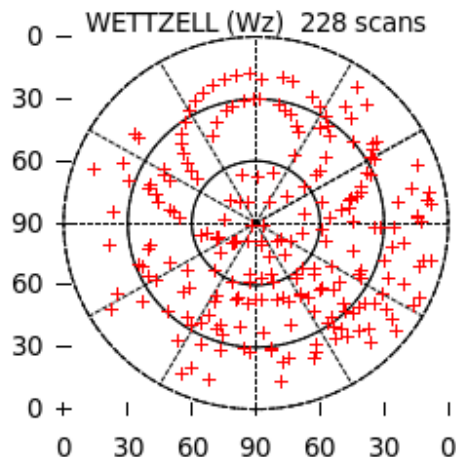
- Analog signals are digitized and time tagged
- Number of correlated bits determine accuracy  
→ large bandwidth
- Data volume of several Terabyte per day
- Data transfer to correlator
  - By network links
  - By disk units
- Hydrogen masers
  - $\Delta f/f \approx 5 \cdot 10^{-15}$
  - Time synchronization  $\leq 1 \mu\text{s}$



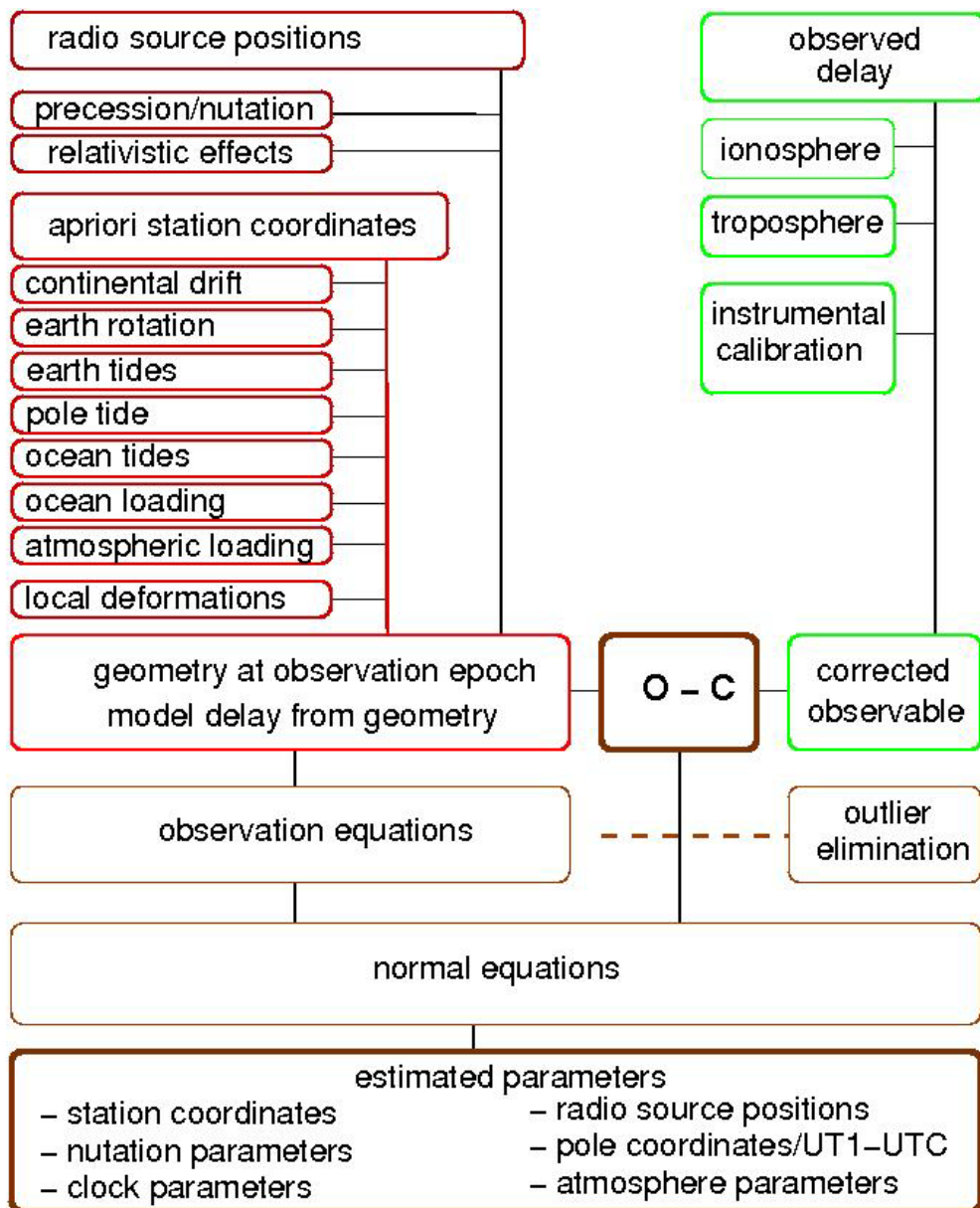




- Recording time 20 – 200 s (= 1 scan)  
(Integration time for signal-to-noise)
- Earth rotates
- Geometry not stationary
- Makes group delay determination complicated
- 1 scan produces one group delay/delay rate

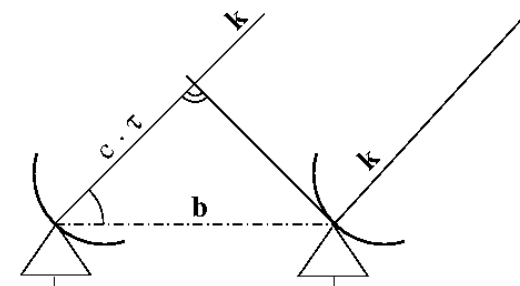


- Multiple scans in one observing session  
(1hr or 24 hour duration)
- Intermediate step:  
**Correlation**



- Input 200 – 400 group delays per baseline per 24 hours

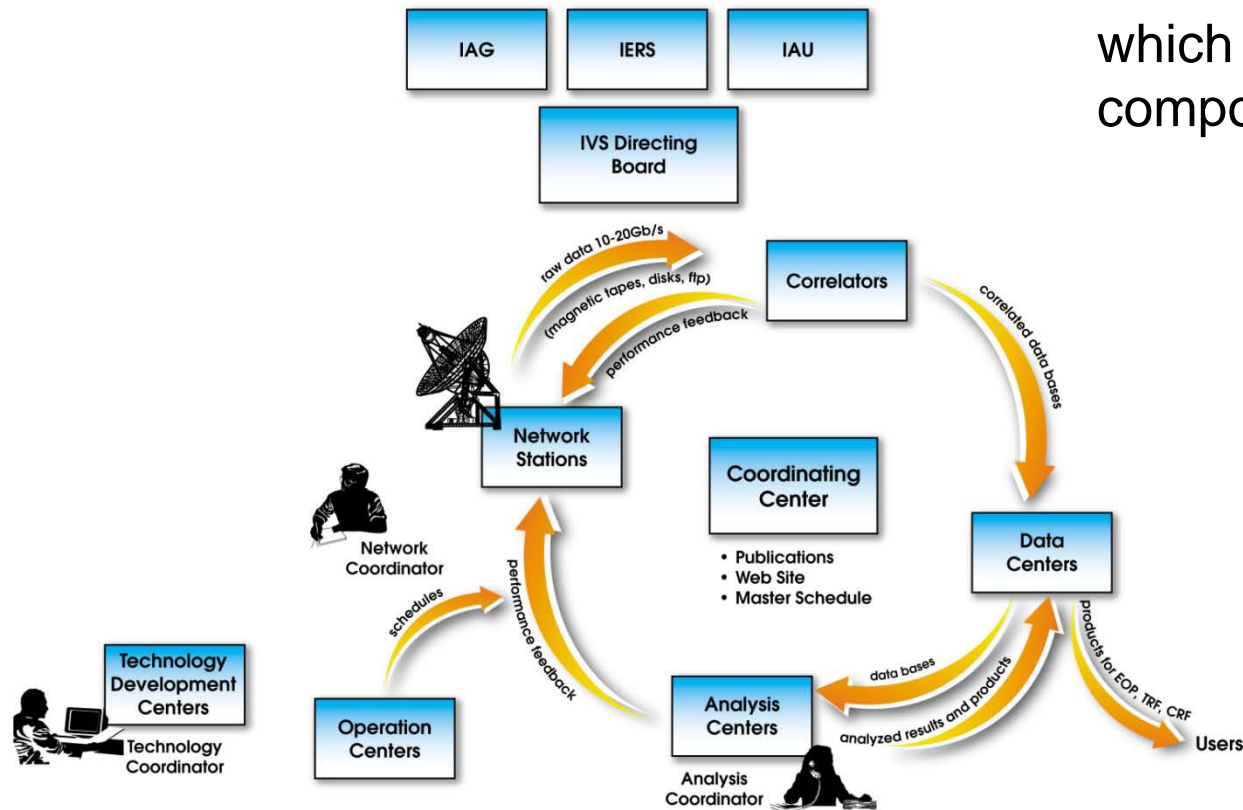
$$\tau = -\frac{1}{c} \vec{b} \cdot R(t) \cdot \vec{k}$$

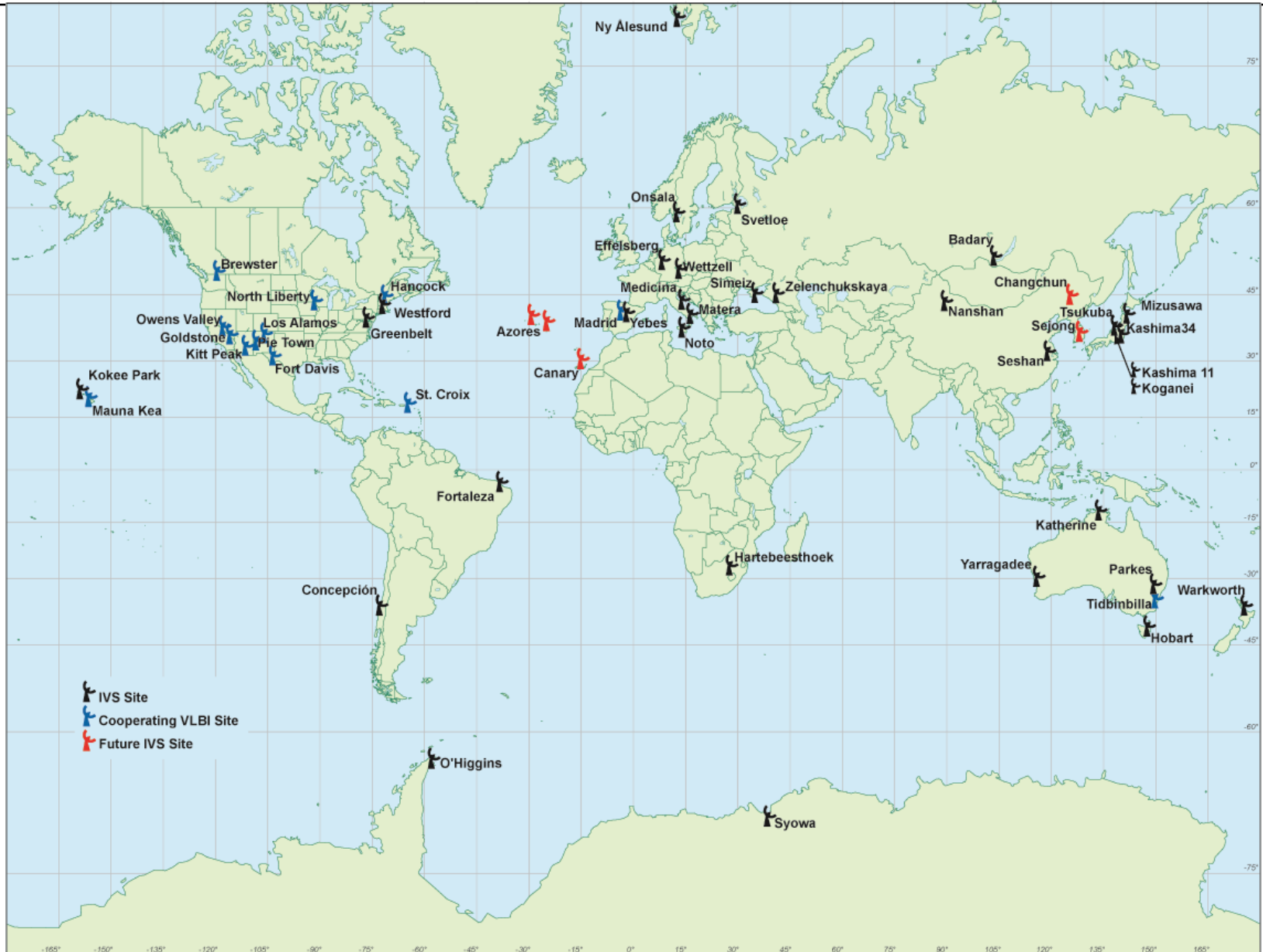


## International VLBI Service for Geodesy and Astrometry

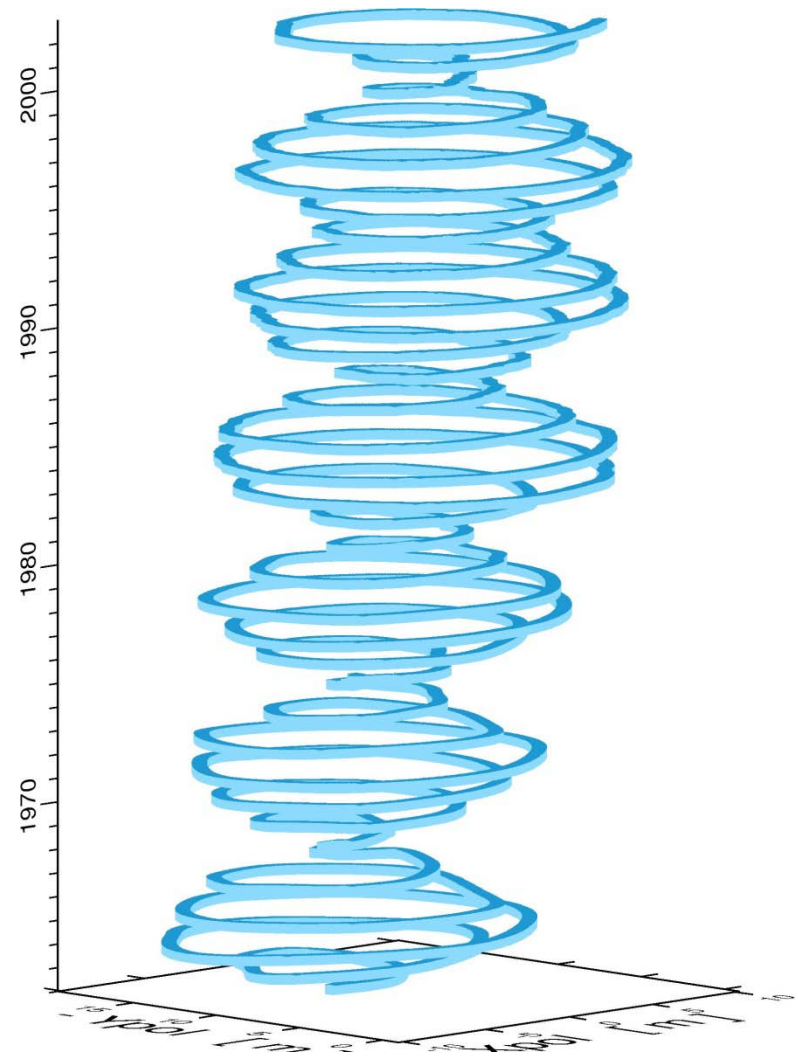
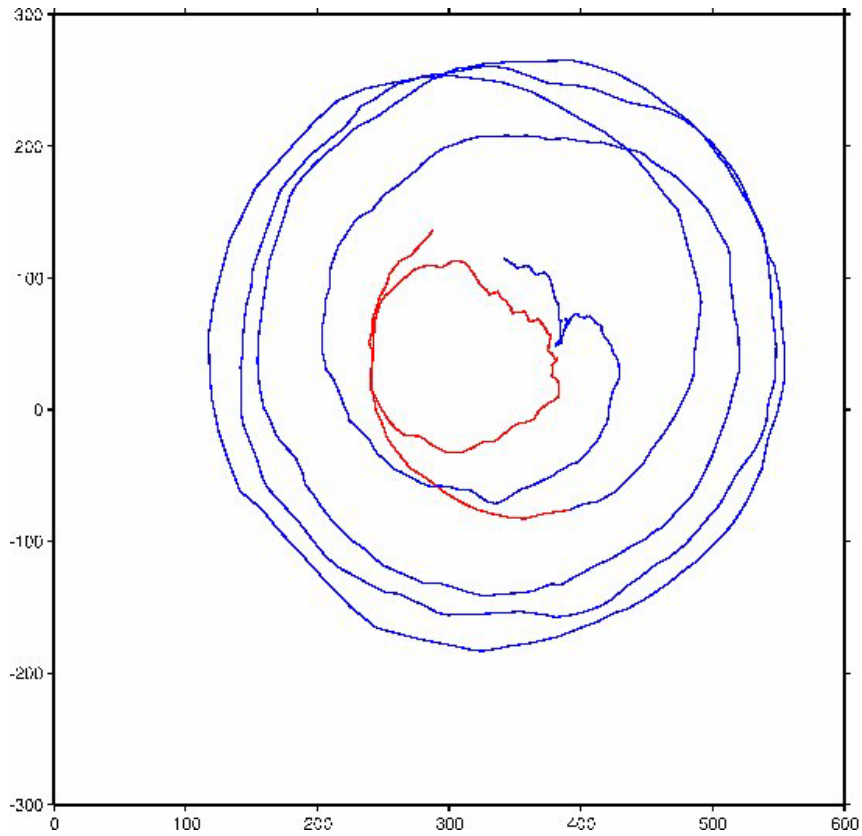
The IVS is an international collaboration of organizations which operate or support VLBI components

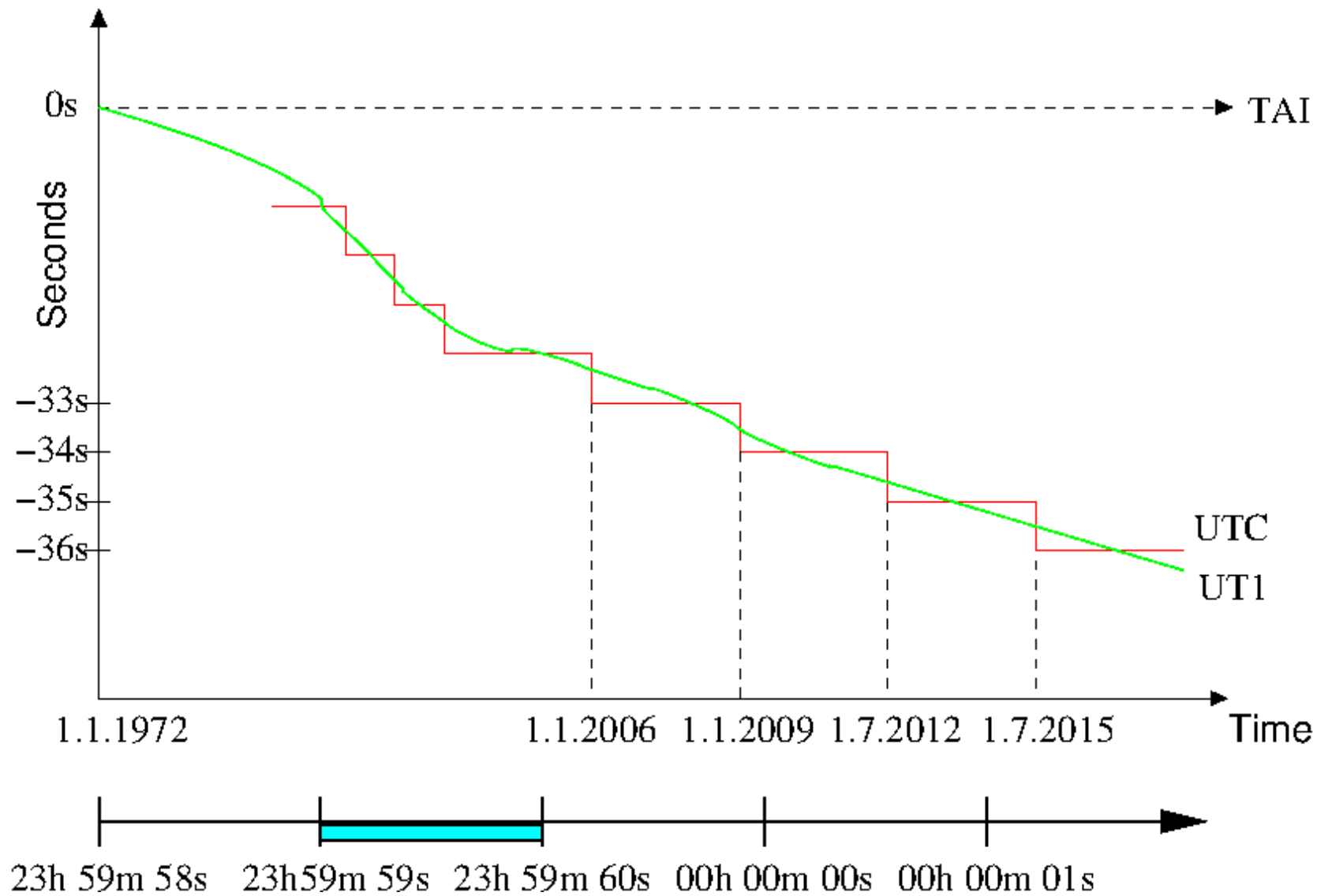
### ORGANIZATION OF INTERNATIONAL VLBI SERVICE

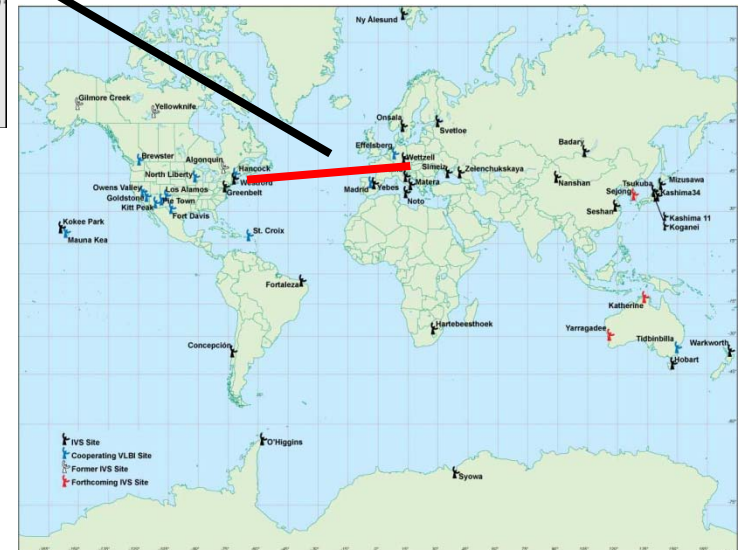
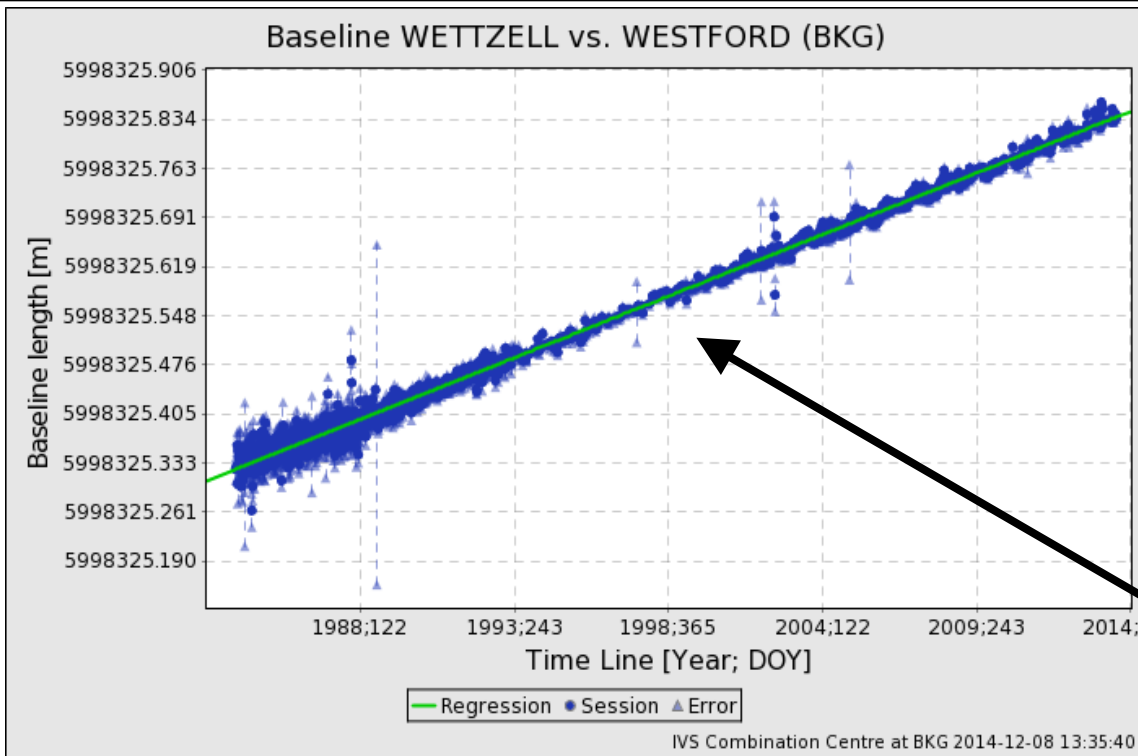






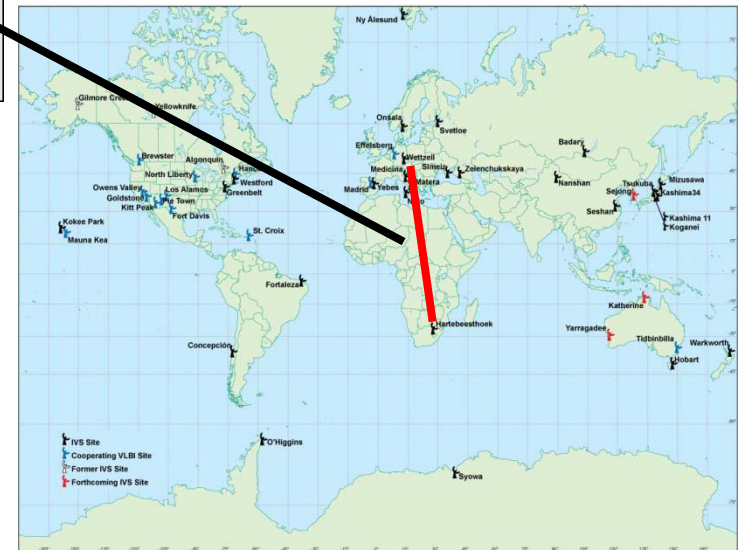
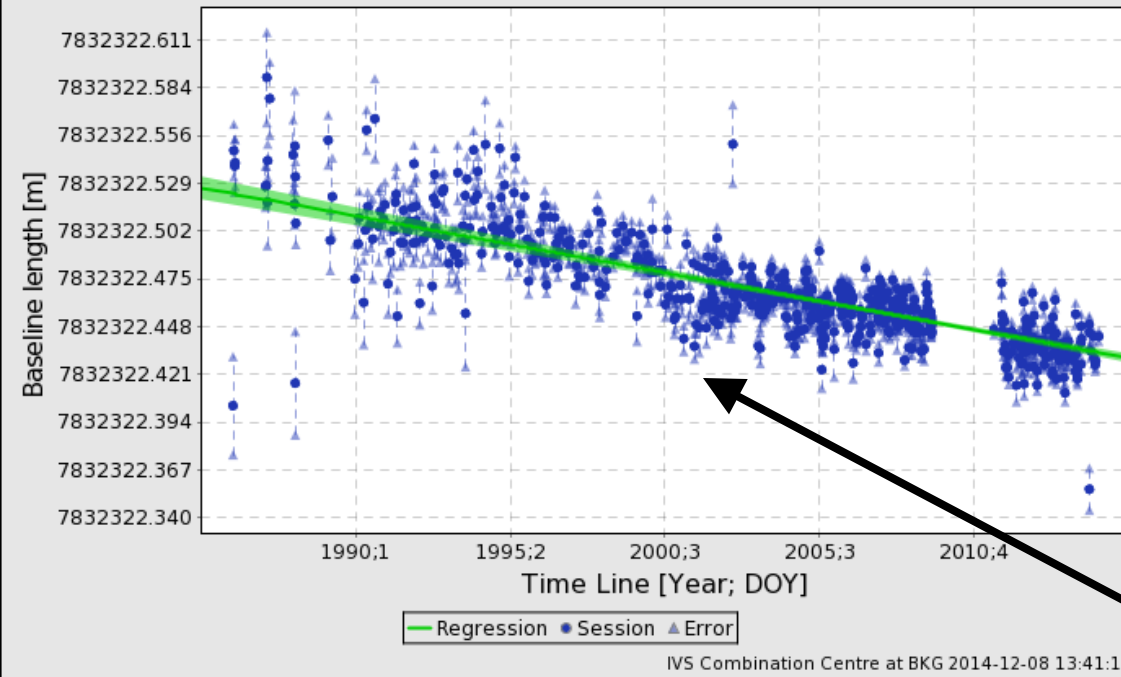




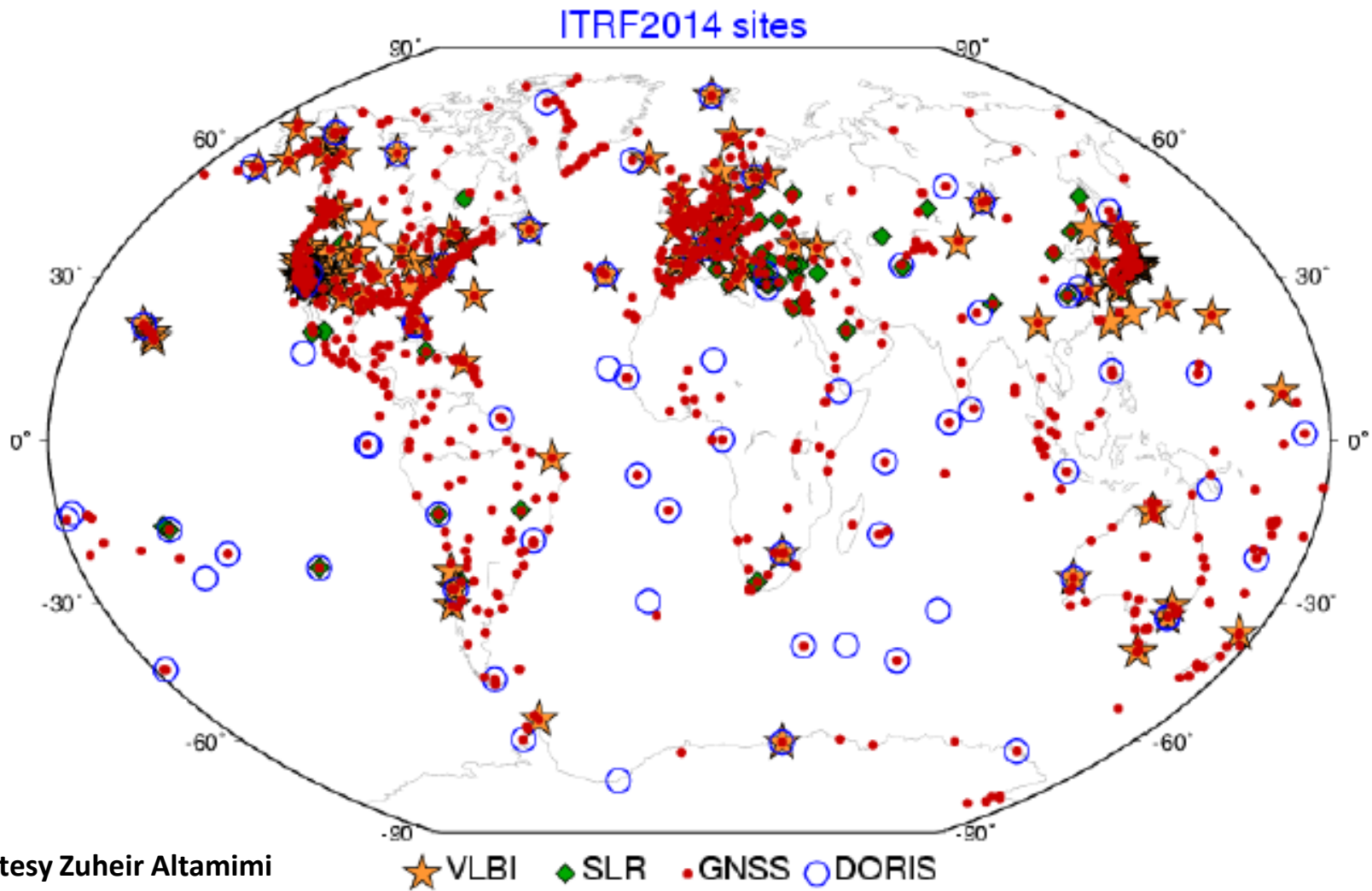




Baseline HARTRAO vs. WETTZELL (GSFC)



## ITRF = International Terrestrial Reference Frame

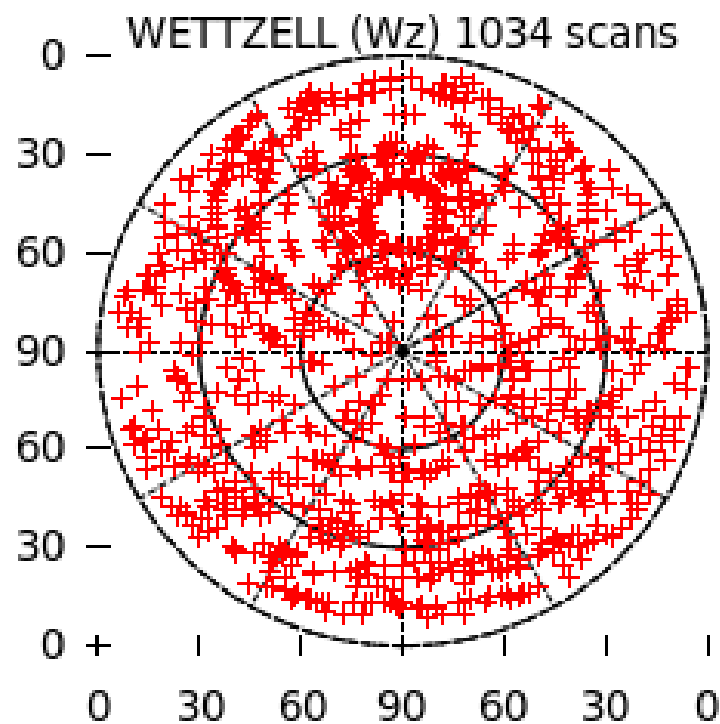
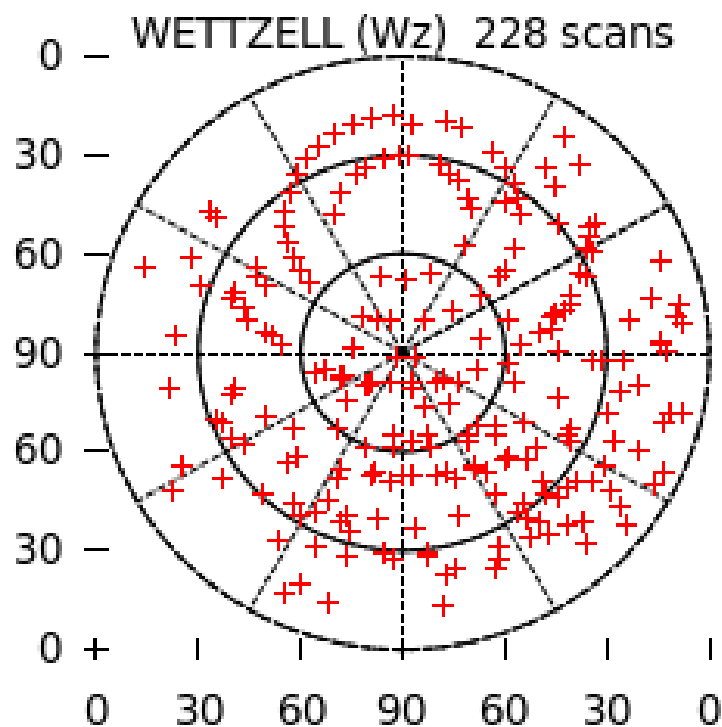


- ITRF2014 STATION POSITIONS AT EPOCH 2010.0 AND VELOCITIES
- VLBI STATIONS

DOMES NB.	SITE NAME	TECH.	ID.	X/Vx	Y/Vy	Z/Vz	Sigmas		
							-----m/m/y-----		
-----									
• 10317S003	Ny-Alesund	VLBI	7331	1202462.5677	252734.4956	6237766.1746	0.0006	0.0006	0.0007
• 10317S003				-.01461	0.00759	0.01039	.00004	.00004	.00004
• 10402S002	Onsala	VLBI	7213	3370605.8407	711917.6739	5349830.8703	0.0007	0.0007	0.0007
• 10402S002				-.01437	0.01454	0.01042	.00004	.00004	.00004
• 10503S002	Metsahovi	VLBI	7385	2892584.8978	1311715.5529	5512640.1298	0.0009	0.0008	0.0012
• 10503S002				-.01643	0.01453	0.01001	.00006	.00006	.00007
• 12338S003	Badary	VLBI	7382	-838200.9878	3865751.5645	4987670.9035	0.0007	0.0008	0.0008
• 12338S003				-.02773	0.00046	-.00274	.00007	.00012	.00016
• 12350S001	Svetloe	VLBI	7380	2730173.7230	1562442.7504	5529969.1152	0.0007	0.0007	0.0007
• 12350S001				-.01863	0.01306	0.00793	.00005	.00005	.00007
• 12351S001	Zelenchukskaya	VLBI	7381	3451207.5991	3060375.3698	4391915.0186	0.0007	0.0007	0.0007
• 12351S001				-.02327	0.01308	0.00644	.00010	.00009	.00013
• 12711S001	Medicina	VLBI	7230	4461369.7463	919597.0658	4449559.3477	0.0010	0.0008	0.0010
• 12711S001				-.01827	0.01879	0.01136	.00005	.00004	.00005
• 12717S001	Noto	VLBI	7547	4934562.8871	1321201.4867	3806484.6874	0.0009	0.0007	0.0008
• 12717S001				-.01786	0.01752	0.01524	.00005	.00005	.00005
• 12734S005	Matera - ASI	VLBI	7243	4641938.5242	1393003.2627	4133325.7273	0.0007	0.0007	0.0007
• 12734S005				-.01886	0.01893	0.01443	.00004	.00004	.00004

- **New generation VLBI infrastructure**

- dense sampling of local sky  
for optimal estimation of atmosphere parameters



**VLBI2010 Global Observing System  
became  
VLBI Global Observing System – VGOS**

**Small and agile telescopes  
Large bandwidth (2 – 14 GHz [3 - 18])  
Flexible frequency allocation  
Dual linear polarization**



- **New generation VLBI infrastructure**

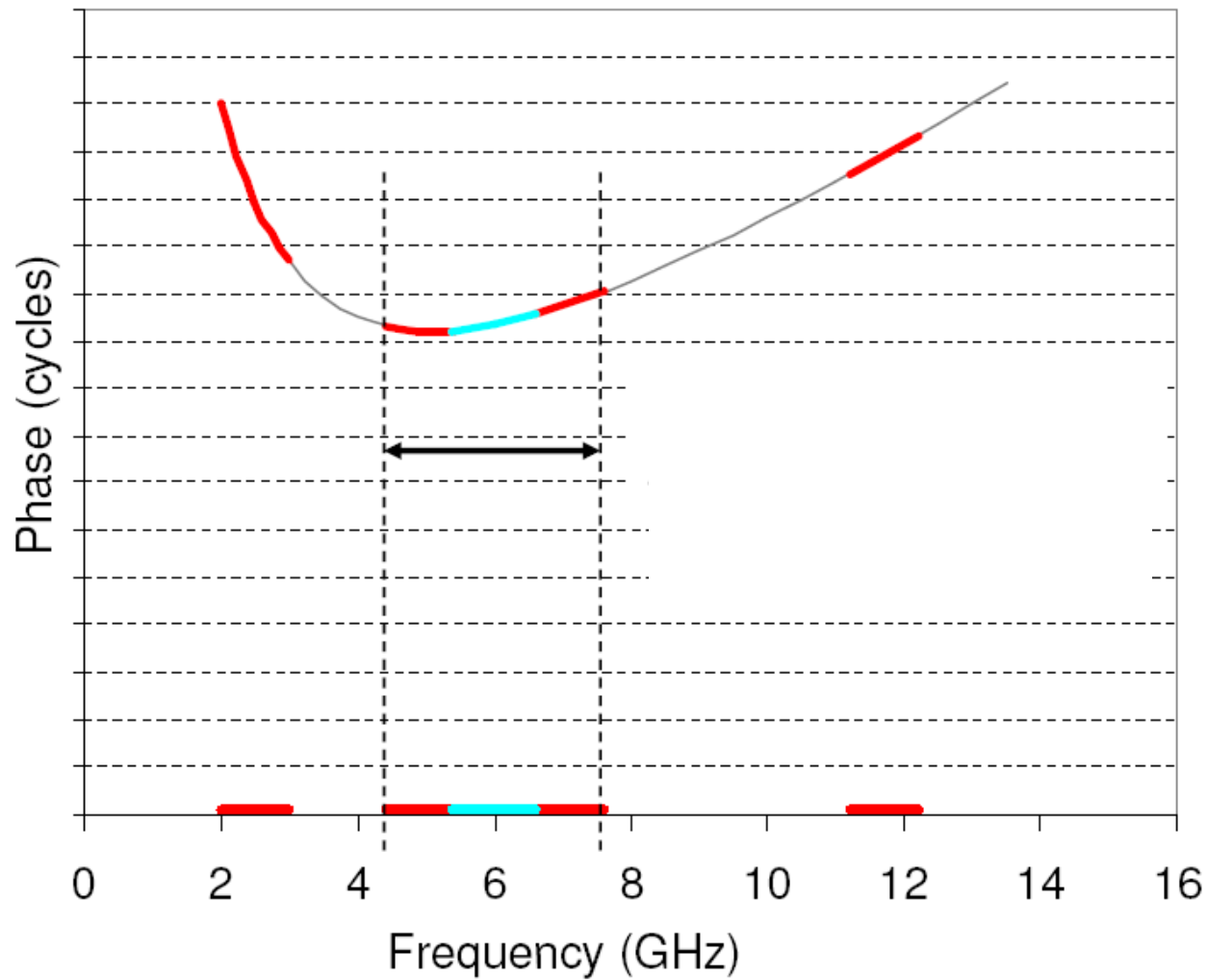
- dense sampling of atmosphere
- agile telescopes
  - small (12 – 13 m)
  - 12°/sec
- up to 2 observations per minute (2880/day)

$$\sigma_{\tau} \propto \sqrt{\frac{1}{A_1 A_2 \cdot B}}$$

**=> Large bandwidth needed**

- wide band receivers (2 – 14 GHz [3 - 18])





Courtesy B. Petrachenko



**Wettzell (DE)**



**Zelenchukskaya (RU)**

Courtesy  
A. Ipatov



**Ishioka (JP)** Courtesy Y. Fukuzaki



**Badary (RU)**

Courtesy  
A. Ipatov



**GGAO (US)**

Courtesy A. Niell



Gómez-González  
et al .(2013)



Yebes (Spain)

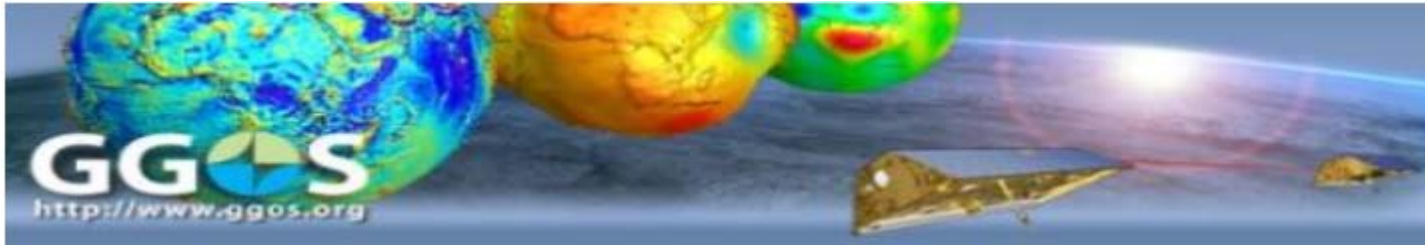
(August 2013) Courtesy: J.A. Lopez



Santa Maria (Eastern Azores)

(Sep. 2014) Courtesy: F. Colomer





## IVS → VGOS → GGOS

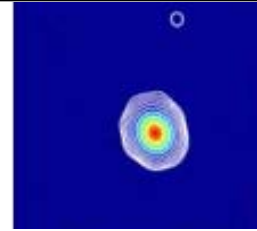
**Global Geodetic Observing System  
of the International Association of Geodesy (IAG)**

### Contribution to GGOS

- **Global distribution** → **Well-designed network**
- **Continuous** → **Economic operations**
- **Stable over decades** → **Monitoring of telescopes and local ties**
- **1 mm/0.1 mm/y** → **Improved technology, better modeling**



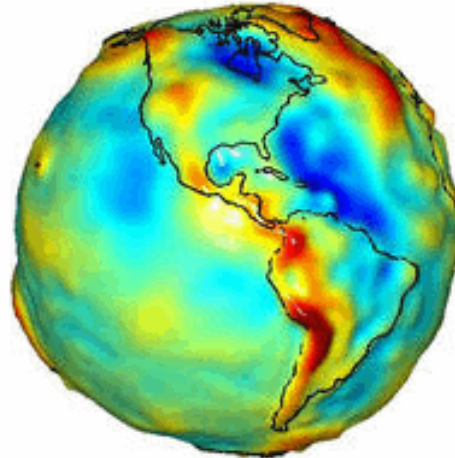
**DORIS**



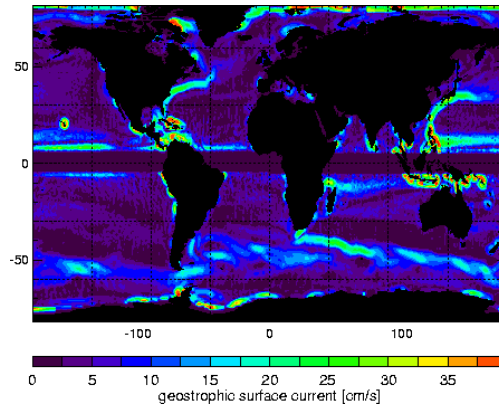
**VLBI**



**GPS**



**SLR**

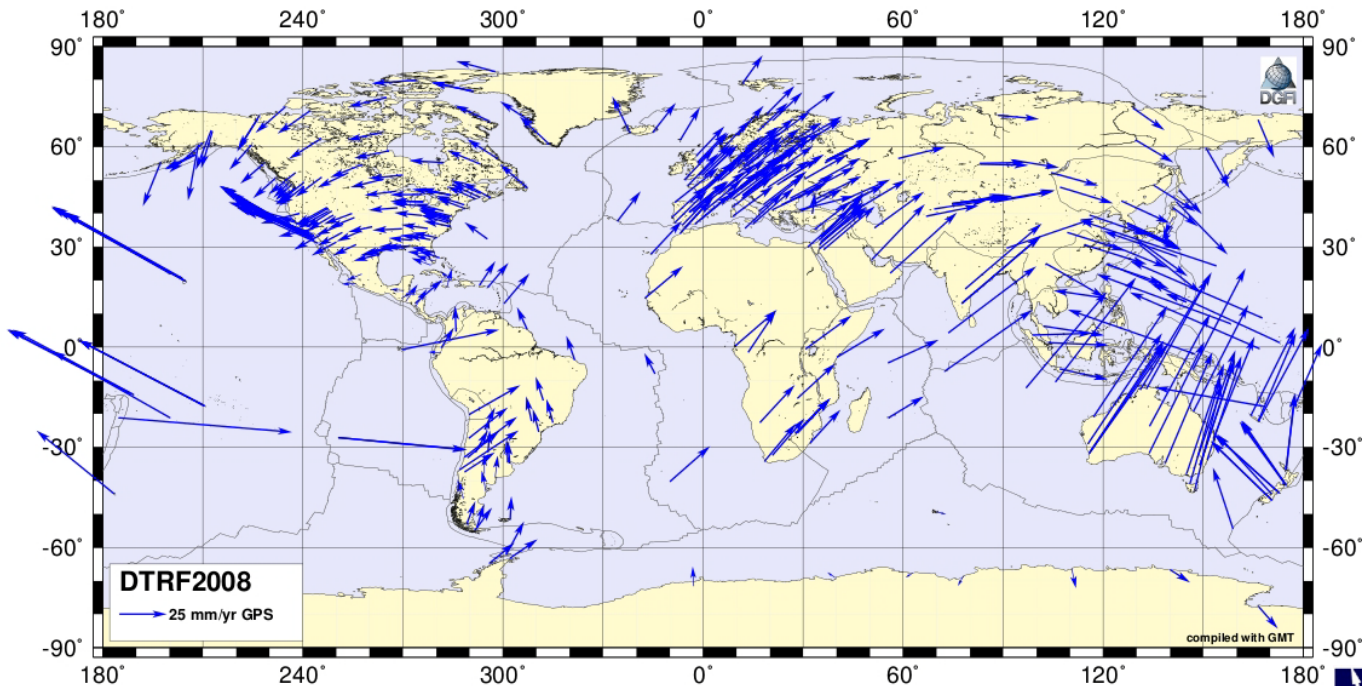




International Association of Geodesy  
Global Geodetic Observing System



## VGOS will be part of GGOS



International  
Association of  
Geodesy